

probe entered the Titan atmosphere providing scientists with the first direct sampling of the moon's chemistry and the first astonishing closeup views of this alien surface as it penetrated the haze that obscured the surface from the earlier Pioneer and Voyager probes sent to Saturn. "The Huygens probe has discovered a new world, Sushil Atreya of the University of Michigan and one of the vast team of scientists studying the rich data sent back by Huygens, said, And there is so much more to learn.

The Aerosol Collector and Pyrolyser (ACP) aboard Huygens collected atmospheric aerosols carried out a preparation involving evaporation, pyrolysis and gas product transfer and then an analysis with gas chromatography-mass spectrometry (GC-MS). The ACP collected two samples one from the top of the descent down to the tropopause (160-40 km) and the second sample in the cloud layer (23-17 km).

The system with its dynamic range of  $10^8$  could identify atmospheric constituents over a mass range from 2 to 146 atomic mass units, and revealed the chemical composition of the photochemical aerosols - in terms of hydrogen, carbon, nitrogen, and oxygen content as well as the relative concentrations of condensates in the lower stratosphere, such as  $C_2$ ,  $H_2$ , H,  $HC_3N$ , HCN, and other small molecules. The relative concentrations of organic condensates within the troposphere, such as methane and ethane and non-condensable constituents, like carbon dioxide, trapped in the collected particles were also determined.

"The GC-MS data provide strong evidence of a thick cloud or haze layer of methane in Titan's middle troposphere around 20 km above the surface," <u>Sushil Atreya</u>

explains, "and there is a reservoir of liquid methane on the surface." Nitrogen is the most abundant gas

Sushil Atreya

in the Titan atmosphere, but Atreya adds that for Titan the role of water on Earth seems to be filled by methane." Heat released by Huygens warmed the surface material beneath the probe and allowed the GC-MS and possibly SSP systems to detect bursts of methane gas as it boiled off from the frozen surface, reinforcing methane's principal role in Titan's geology and atmospheric meteorology forming clouds and precipitation that erodes and abrades the surface.

Analytical evidence of such an abundance of methane is providing important clues that might answer the perplexing question of why there is so much methane present in the first place. "The big question for Titan is how the methane gets replenished," Atreya explains, "In the absence of



The Aerosol Collector and Pyrolyser (ACP). (Image credit: Service d'Aeronomie, Centre National de la Research Scientifique, CNRS)

recycling it would be destroyed by the Sun's ultraviolet light in 10 million years, which would in turn lead to a gradual collapse of Titan's atmosphere." The GC-MS measurements, together with supporting data from Huygens' other instruments, suggest that recycling does occur and Atreya is confident that the ongoing data analysis will provide a clear picture of Titan's hazy atmosphere in the coming months.

Some observers have speculated that because the main means of methane recycling in Earth's atmosphere is due to the organic activity of peat bogs, rice fields and ruminant animals, that perhaps Titan harbours some kind of alien life form that continuously replenishes the gas destroyed by sunlight's photo-oxidative effects. But, those hoping for ET have to face the cold facts. Titan is not a pleasant place for life it is much too cold for liquid water at -180 Celsius. A more plausible explanation may lie in vast oceans of methane that lie on or beneath the surface of Titan that are constantly feeding the atmosphere.

For more on the Cassini-Huygens Mission, read the special issue of "spectral lines" on our sister site, spectroscopyNOW.com



## **Related links:**

- Cassini-Huygens homepage, NASA
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- <u>Gas</u> Chromatograph and <u>Mass</u> <u>Spectrometer Instrument (GCMS)</u>, ESA website
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Article by David Bradley

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